

# Space time analysis of extreme values <sup>a</sup>

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4th EVA conference, Gothenburg, August 15-19 2005

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<sup>a</sup>Joint paper with Bruno Sanso, University of California, Santa Cruz, U.S.A

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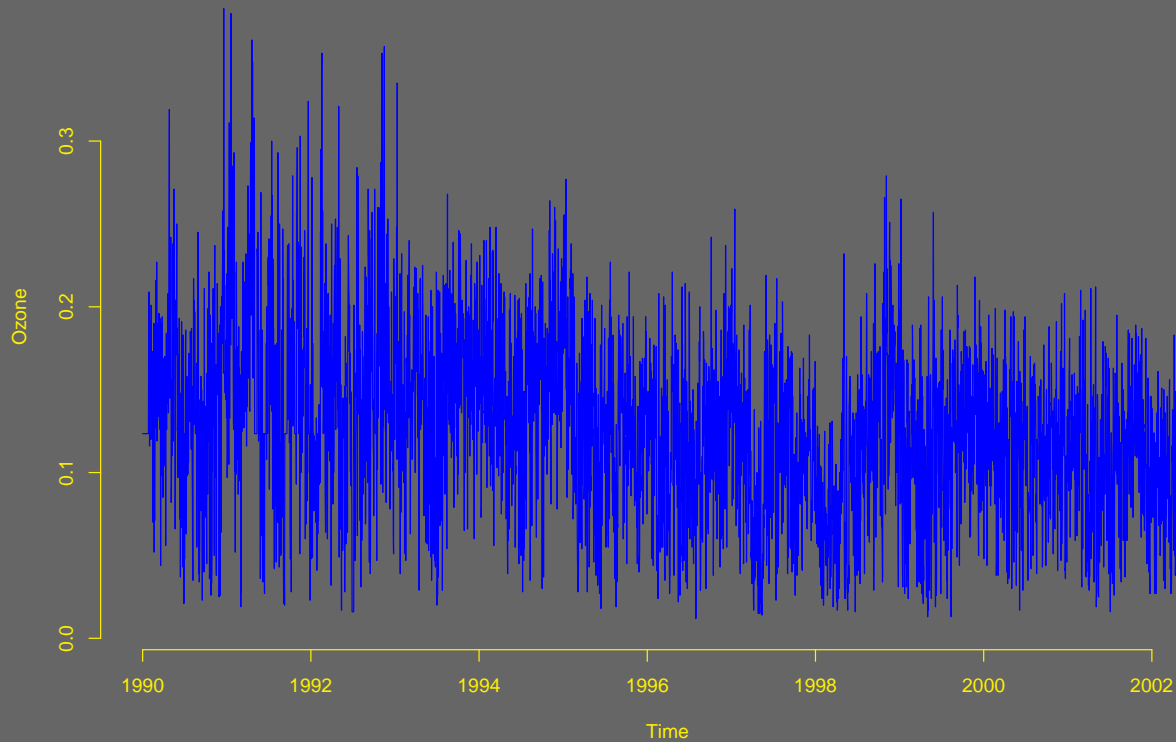
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- Extreme values of rainfall in Venezuela.

Figure 1: Daily maximum values of ozone levels.



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- $+$  denotes the positive part of the argument.
- $\xi > 0$  *Fréchet* family;  $\xi < 0$  the *Weibull* family;  
 $\xi \rightarrow 0$  *Gumbel* family.

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- These ideas had been developed into a Bayesian hierarchical modeling framework. Smith et al. (1997); Assuncao et al. (2004); Casson and Coles (1999); Gilleland et. al. (2004).

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- Non-stationarity can also be included for the shape and/or scale parameters:  $\sigma_t = \exp(\beta_0 + \beta_1 t)$ ;  $\xi_t = \beta_0 + \beta_1 t$  or  $\xi_t = \beta_0 + \beta_1 t + \beta_2 t^2$ .

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- We propose the use of Dynamic Linear Models (DLM) as in West and Harrison (1997) to model the parameter changes in time.

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- For  $z_1, z_2, \dots, z_m$ ,  $z_t \sim GEV(\mu_t, \sigma, \xi)$

$$H_t(z_t) = \exp \left\{ -[1 + \xi(z_t - \mu_t)/\sigma]_+^{-1/\xi} \right\}$$

$$\mu_t = \theta_t + \epsilon_t; \quad \epsilon_t \sim N(0, V)$$

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- $\pi(\sigma) \sim LN(m_\sigma, s_\sigma)$ ;  $\pi(\xi) \sim N(m_\xi, s_\xi)$ .
- $\theta_0 \sim N(m_0, C_0)$ ;  $V \sim IG(\alpha_v, \beta_v)$ ;  $\tau \sim IG(\alpha_\tau, \beta_\tau)$

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- $\mu_t$  follows a *first order polynomial* DLM with state vector  $\theta_t$ .

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$$\begin{aligned}\mu_t &= F_t' \theta_t + \epsilon_t; \quad \epsilon_t \sim N(0, V) \\ \theta_t &= G_t \theta_{t-1} + \omega_t; \quad \omega_t \sim N(0, W)\end{aligned}$$

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- $\theta_t$  is a  $k \times 1$  state vector;
- $F_t$  is a  $k \times 1$  regressor vector;
- $G_t$  is a  $k \times k$  evolution matrix;
- $V$  is an observational variance and
- $W$  is a  $k \times k$  evolution covariance matrix.

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- For  $\theta_t$ , we apply *Forward Filtering Backward Simulation* (FFBS) as in Carter and Kohn or Frühwirth-Schnatter (1994).

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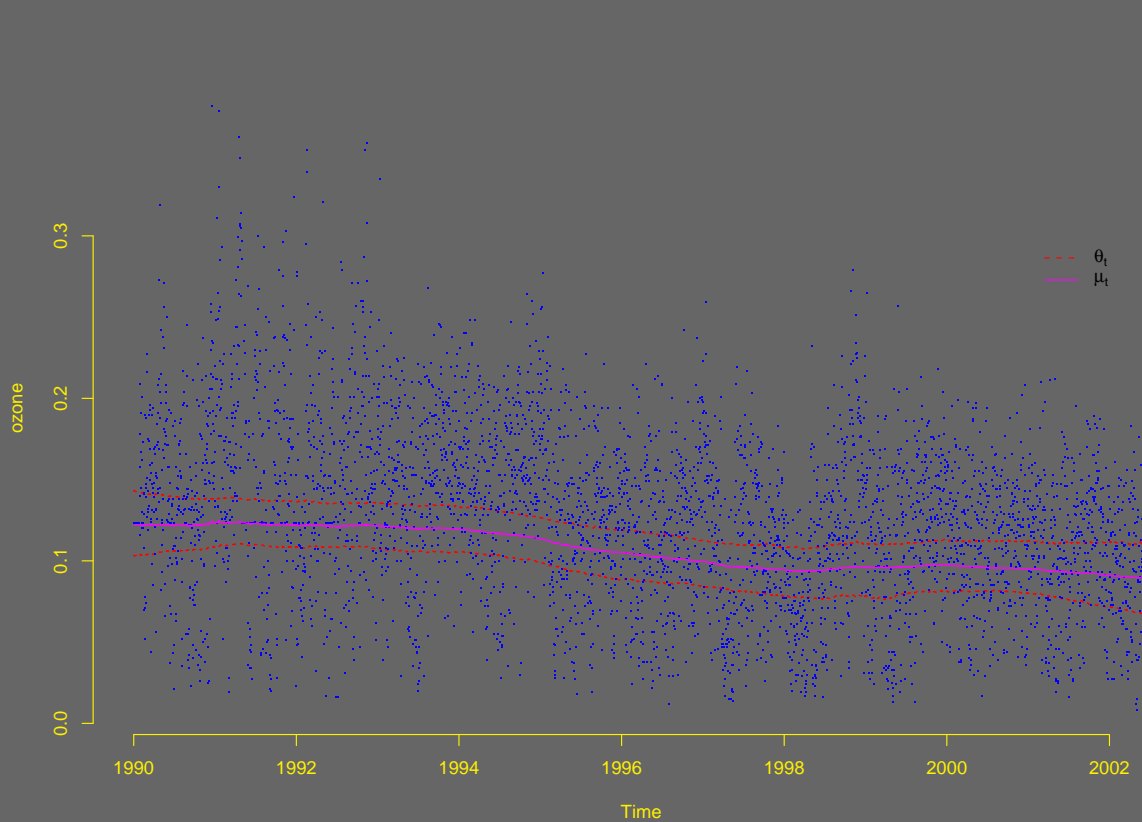
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  - Dynamics for scale/shape parameters. Sequential updating with *Particle Filters*.
  - Observation variance equals zero.
  - No space or space-time structure.

Figure 2: Posterior mean for  $\mu_t$ , and 90% probability interval for  $\theta_t$ ; 1990-2002 data



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$$\begin{aligned}z_t &\sim GEV(\mu_t, \sigma, \xi) \\ \mu_t &= \theta_t + \beta(t - \bar{t}) + \epsilon_t; \quad \epsilon_t \sim N(0, V) \\ \theta_t &= \theta_{t-1} + \omega_t; \quad \omega_t \sim N(0, \tau V)\end{aligned}$$

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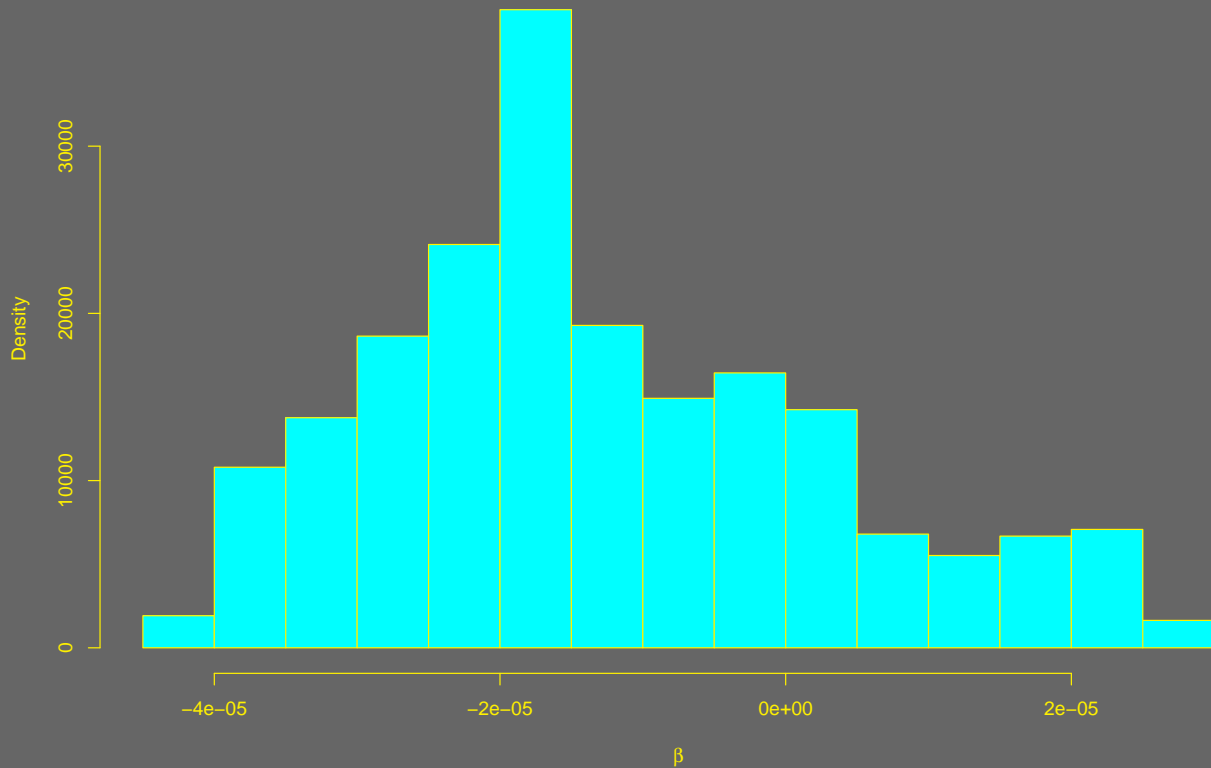
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  - This defines a *Gibbs sampler* scheme that produces posterior samples for  $\beta$  and  $\theta_t$ .
- In fact,  $Pr(\beta < 0|Z) \approx 0.79$  indication of a decreasing trend.

Figure 3: Posterior distribution for  $\beta$ ; ozone data 1990-2002.



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- The model is:

$$Y_t \sim GEV(\mu_t, \sigma, \xi)$$

$$\mu_t = \theta_t + \beta_t X_t + \epsilon_t$$

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Figure 4: Maxima monthly rainfall values in Venezuela and NAO index

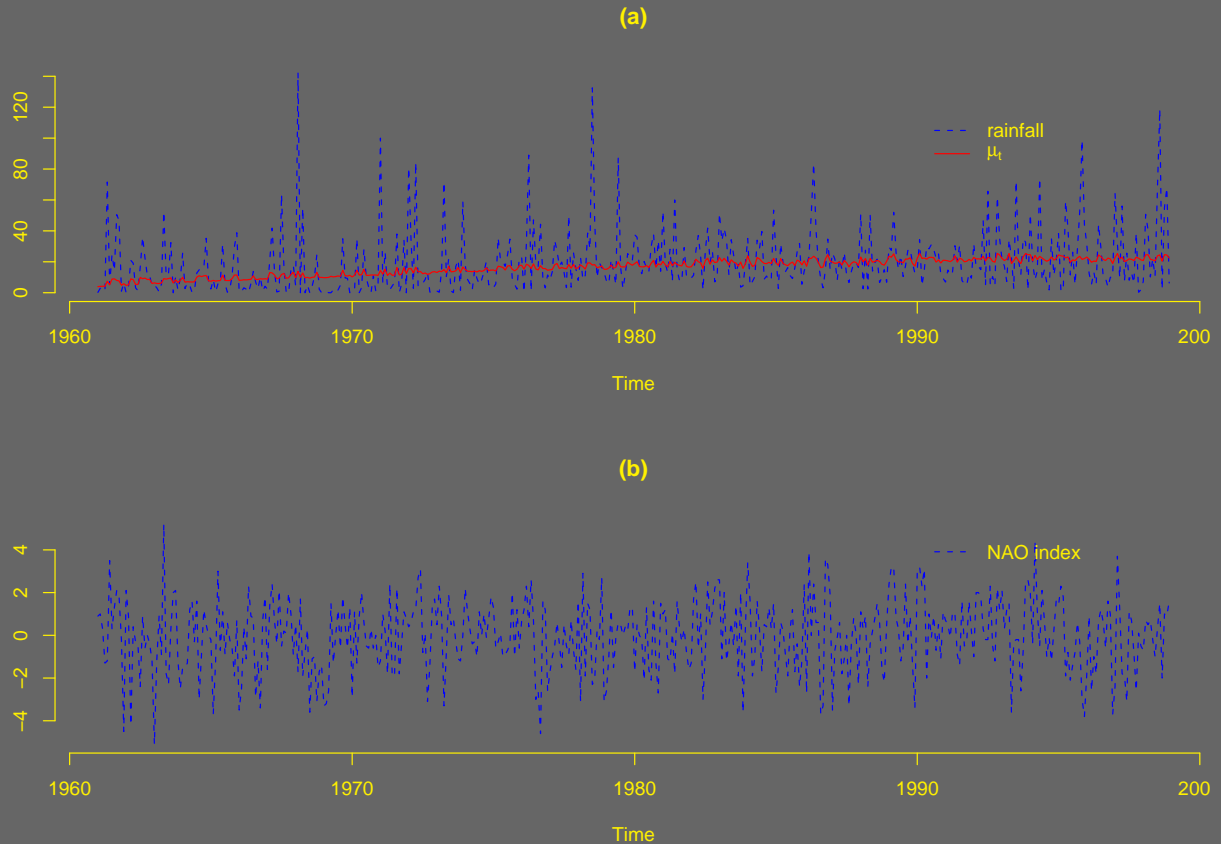
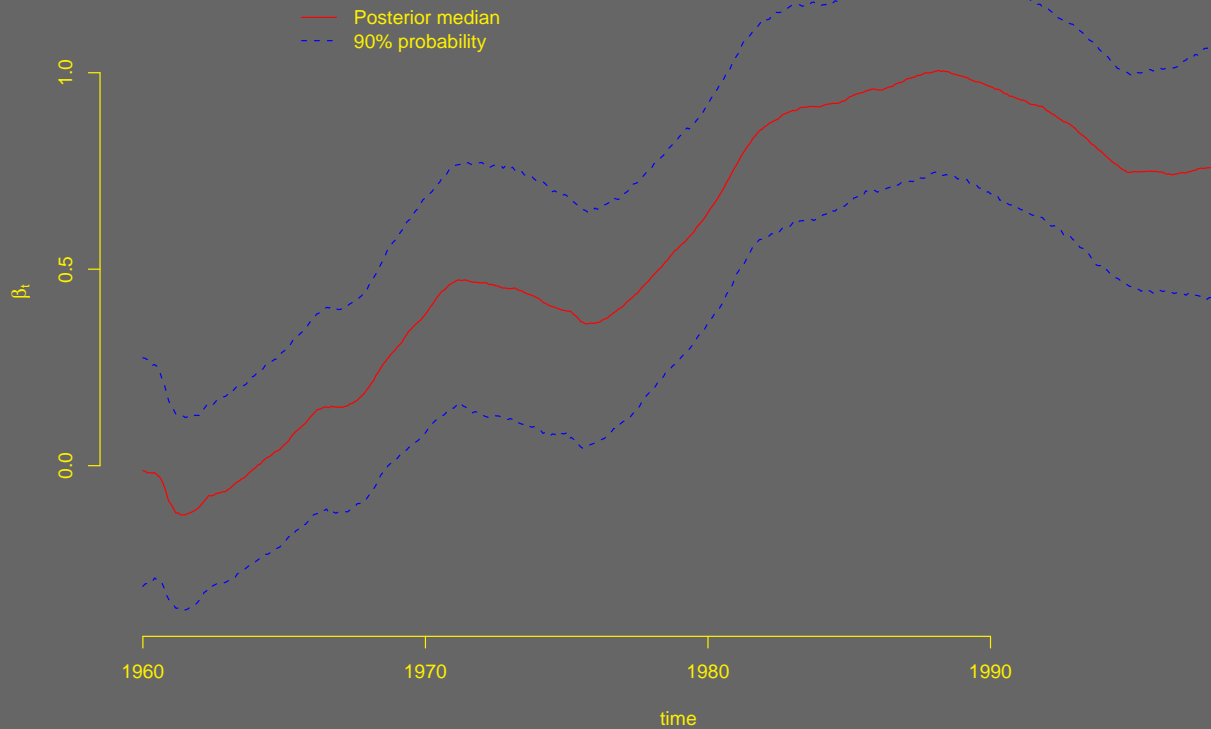


Figure 5: Posterior median and 90% probability intervals for  $\beta_t$



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$K^t$  is a  $n_t \times \kappa$  matrix given by

$$K_{ij}^t = k(s_i - \omega_j), \quad t = 1, \dots, m$$

$$\epsilon_t \sim N(0, \sigma_\epsilon^2), \quad t = 1, \dots, m$$

$$\nu_t \sim N(0, \sigma_\nu^2), \quad t = 1, \dots, m$$

$$x_1 \sim N(0, \sigma_x^2 \mathbf{I}_\kappa)$$

- $k(\cdot - \omega_j)$  defines a smoothing kernel.
- $\omega_1, \dots, \omega_\kappa$  are spatial sites where kernels are centered.
- $x_t$  is interpreted as a *latent process*

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- Possible kernels are:
  - *Gaussian*:  $k(s) \propto \exp \{-\|s\|^2/2\eta\}$  ( $\eta > 0$ ).
  - *Exponential*:  $k(s) \propto \exp \{-\|s\|/\eta\}$  ( $\eta > 0$ ).
  - *Spherical*:  $k(s) \propto \left(1 - \frac{\|s\|^3}{r^3}\right)^3 I[s \leq r]$ .

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- We define a DLM on  $\mu_t$ :

$$\mu_t = K' \theta_t + \epsilon_t;$$

$$\theta_t = \theta_{t-1} + \nu_t$$

- $\theta_t = (\theta_{t,1}, \dots, \theta_{t,\kappa})'$ ,  $\epsilon_t = (\epsilon_{t,1}, \dots, \epsilon_{t,\kappa})'$ ,  $\nu_t = (\nu_{t,1}, \dots, \nu_{t,\kappa})'$ .

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- With a Gaussian kernel,  $K'$  is an  $S \times \kappa$  matrix with entries:

$$K'_{ij} = K(s_i - \omega_j);$$

$$K(s_i - \omega_j) \propto \exp(-d \|s_i - \omega_j\|^2 / 2)$$

- $s_i$  is the position of station  $i$ .
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- $d$  is a range parameter;  $d = c\phi$ ;  $1/2 < c < 2$ ;  $\phi = \text{knot distance}$ .

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- The log-likelihood is equal to

$$\begin{aligned}
 l(\boldsymbol{\theta}) = & -mS \log \sigma - \sum_{t=1}^m \sum_{s=1}^S \left[ 1 + \xi \left( \frac{z_{s,t} - \mu_{s,t}}{\sigma} \right) \right]_+^{-1/\xi} \\
 & - \left( 1 + \frac{1}{\sigma} \right) \sum_{t=1}^m \sum_{s=1}^S \log \left[ 1 + \xi \left( \frac{y_{s,t} - \mu_{s,t}}{\sigma} \right) \right]_+
 \end{aligned}$$

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- The range parameter  $d$  is assumed fixed,  $d = c\phi$ .

Figure 6: RAMA stations, kernel and interpolation grid positions.

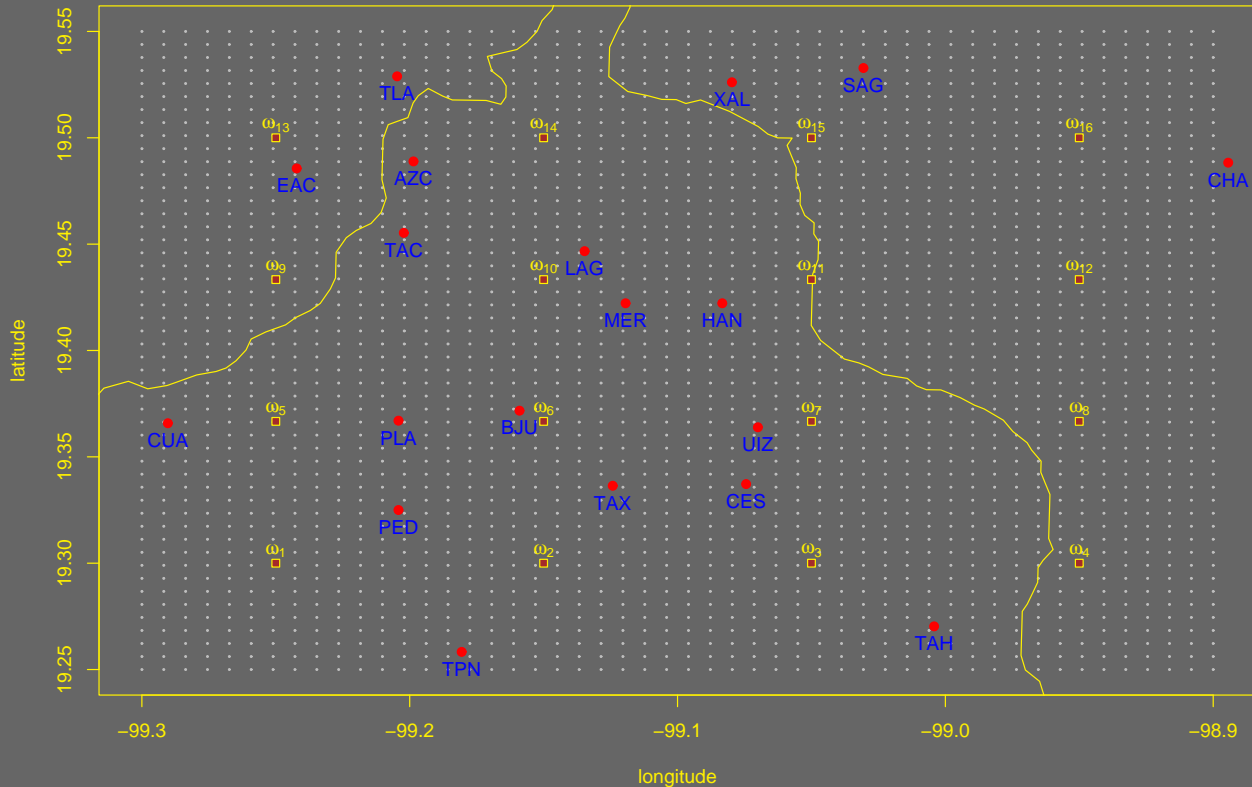


Figure 7: Daily maxima for 1999 and posterior estimates of 0.5 quantile.

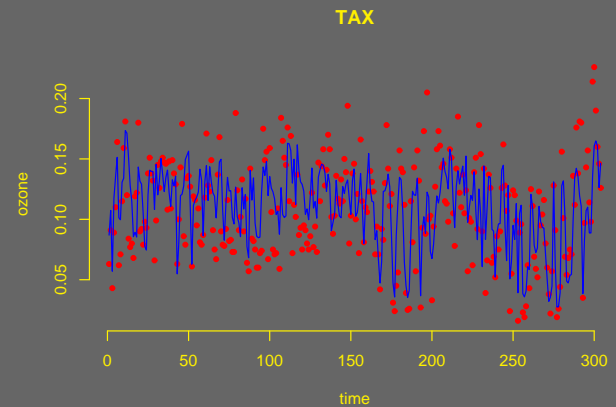
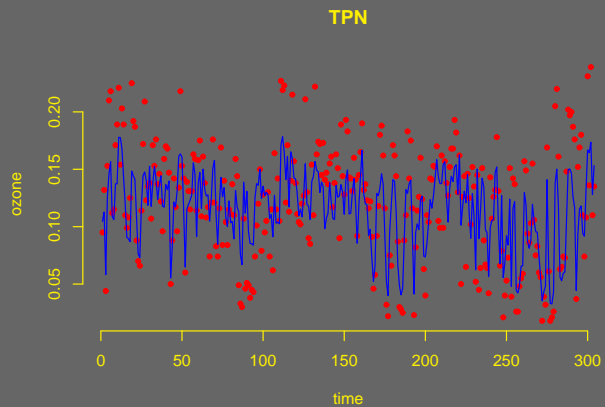
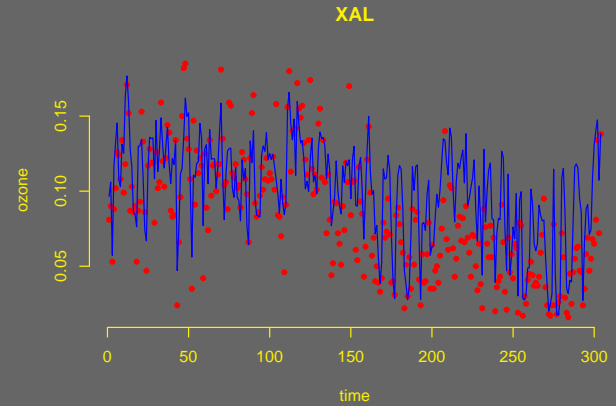
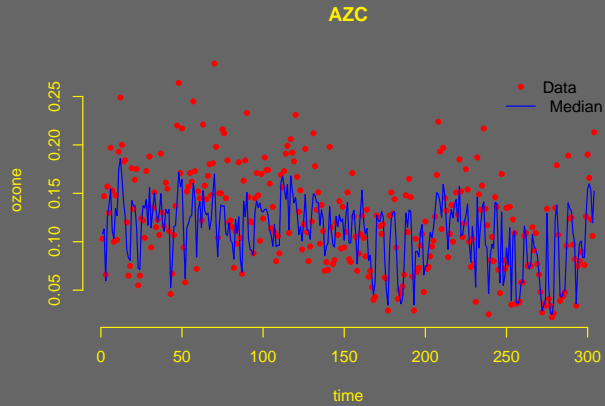


Figure 8: Posterior estimate of the 0.5 quantile of the space-time GEV distribution for a  $50 \times 50$  resolution grid.

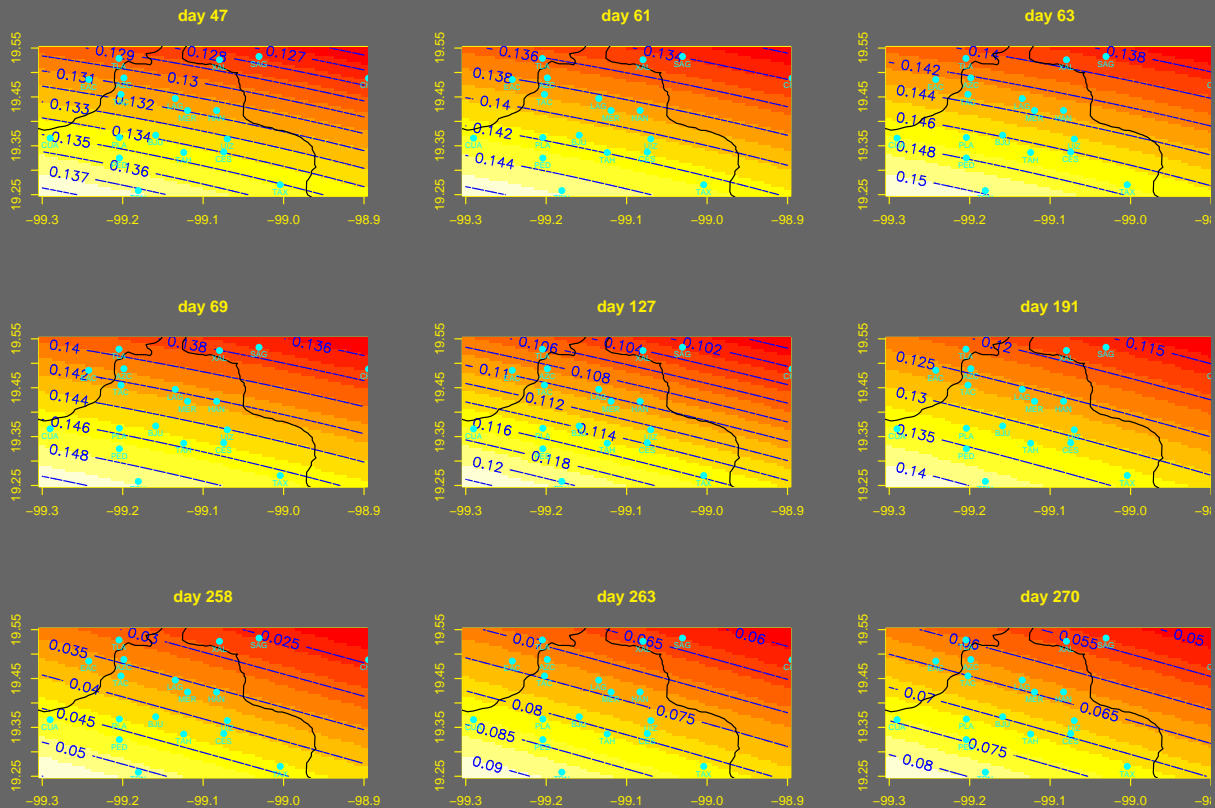
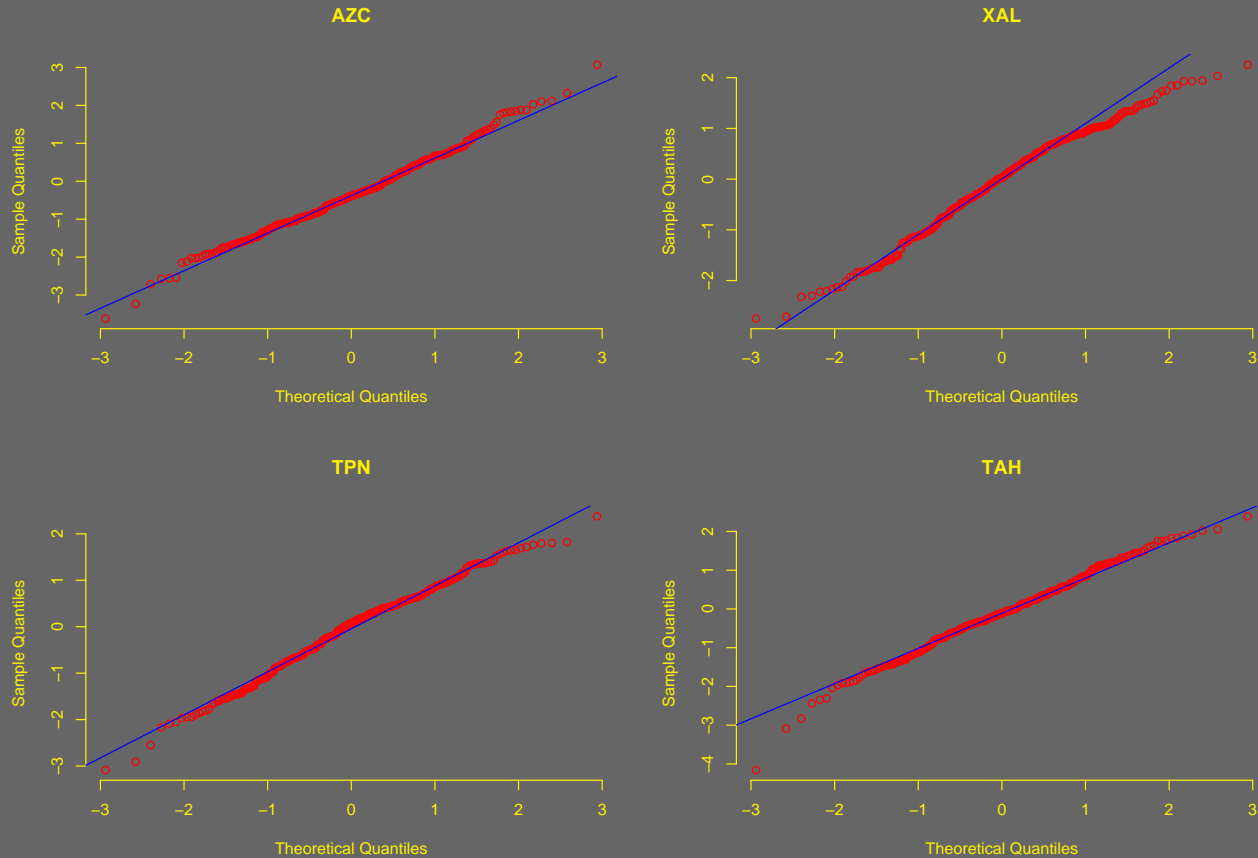


Figure 9:  $u_t = G(y_t)$  diagnostics based on leaving one station out



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